

Motivation

- Clouds in hot Jupiter atmospheres form through the condensation of gas species on small solid particles called cloud condensation nuclei (CCN)
- ► These CCN are created from the gas phacse through nucleation of molecular clusters of species like TiO_2
- ► To accurately calculate nucleation rates, we need to accurately determine thermochemical properties of these molecular clusters.
- GOAL: Find the minimum energy cluster geometry and thermochemical properties for TiO₂ molecular clusters of size N = 3 - 15.

Methods



- We employ a funnel approach to explore the geometry parameter-space for cluster candidates
- ► Three different levels of approximation are employed to calculate binding energies:
- ▷ Force field
- Density-Functional based tight binding (DFTB)
- Density functional theory (DFT)
- **Steps in the Funnel**
- 1. Determine the best parameterisation for DFT calculations from experimental results for the monomer
- 2. Optimise known small cluster isomers with DFT parameterisation
- 3. Use optimised small cluster isomers to calibrate force-field and DFTB parameterisations (dashed arrows)
- . Generate candidate geometries
- 5. Optimise candidate geometries to find a global minimum

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Summary and Outlook

- impacting cloud formation
- efficient at lower temperatures

Nucleation rates



- homogeneous, homo-molecular nucleation

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